

**PROPOSAL COVERSHEET**

DEADLINES: 1st of Feb., June, Oct.  
 EMAIL TO: toosoc@nrao.edu

rcvd:
-------

- (1) Date Prepared: August 29, 2007  
 (2) Title of Proposal: 3C 111 in Outburst (Target of Opportunity)

(3) AUTHORS (Add * for new location)	INSTITUTION	E-mail	Students Only		
			G/U	For Thesis?	Ph.D. Year
M. Kadler	NASA GSFC	Matthias.Kadler@nasa.gov			
T.P. Krichbaum, E. Ros, Y.Y. Kovalev, L. Fuhrmann, M. Perucho	MPIfR	tkrichbaum, ros, ykovalev, lfuhrmann, perucho@mpifr- bonn.mpg.de			
I. Agudo	IAA-CSIC	iagudo@iaa.es			
A.P. Marscher, S.G. Jorstad	Boston Univ.	marscher, jorstad@bu.edu			

(4) Related previous or current VLBI proposal(s): GL029A  Resubmission

(5) Contact author for scheduling: Matthias Kadler (6) Telephone: +1 301-286-4390  
 Address: Astrophysics Science Div., Code 662 NASA GSFC  
 Greenbelt, MD 20771, USA Fax: +1 301-286-1684

(7) Scientific Category:  astrometry & geodesy  galactic  extragalactic  other:

Rapid Response Science:  Known Transient  Exploratory  Target of Opportunity

(8) Wavelength(s) requested (those not available on the global network are indicated with a small circle):

- 90cm  50cm  30cm  21cm  18cm  13cm  6cm  5cm  3.6cm  3.6/13cm  
 2cm  1.3cm  7mm  3mm  
 Global Network standard bands  Special frequencies: \_\_\_\_\_

(9) Recording format:  Default continuum setup (VLBA only),  VLBA/MkIV,  MkIII: Mode \_\_\_\_\_

Bandwidth per BaseBand channel: \_\_\_\_\_

Aggregate bit rate: 512 ( 16 BB channels at 16 MSamples/sec of  1 bit,  2 bit )

(10)  Multi-epoch observation: \_\_\_\_\_ epochs of \_\_\_\_\_ hours each, separated by \_\_\_\_\_

(11) Network	Requested antennas	Total time requested
EVN & MERLIN		
VLBA		14 hr
other NRAO		
GMVA	Effelsberg, Pico Veleta, Pl. de Bure, Onsala, Metsahovi, + VLBA (all available)	14 hr
Non-VLBI Instruments		

(12) ABSTRACT (Do not write outside this space. Please type)

Millimeter flux density measurements performed at IRAM and the SMA reveal an ongoing major flux-density outburst in the radio galaxy 3C 111. Recent measurements with the IRAM 30-m telescope show more than 8 Jy at 1 mm and more than 12 Jy at 3 mm! We would like to use this unique opportunity and request 14 hours of GMVA time to observe the very early state of the flaring jet base. We will study an expected forward-reverse shock-in-jet structure in response to an enhanced injection of relativistic plasma in the 3C 111 jet utilizing the unprecedented angular resolution and the high observing frequency of the GMVA.

- (13) Observation type:  Interferometry,  Spectroscopy,  Pulsar,  Phase referencing
- (14) Proposal is  Suitable  Unsuitable for dynamic scheduling.
- (15) Polarization:  Single Polarization  Dual Circular Polarization  
 Global network standard for single polarization is LCP for all  $\lambda$ s except 13cm (RCP) and 3.6cm (RCP).
- (16) Tape usage (Show <recording time>/<total time>): 0.49
- (17) Assistance required:  
 Observation Setup:  Consultation,  Extensive help,  Observe file preparation  
 Postprocessing:  Consultation,  Extensive help,  Calibration service
- (18) Processor:  Socorro,  JIVE,  Haystack,  Bonn,  Washington,  Other \_\_\_\_\_  
 Special processing:  XPol,  Pulsar gate,  Multiple Fields: \_\_\_\_\_  
 Averaging time: \_\_\_\_\_ Spectral channels per baseband channel: \_\_\_\_\_  
 Other special processing: \_\_\_\_\_
- (19) Postprocessing Location: MPIfR, NASA GSFC
- (20) Source list:  J2000  B1950  
 If more than 4 sources, please attach list. If more than 30, give only selection criteria and GST range(s)

	Source 1	Source 2	Source 3	Source 4
Name(s)	3C111			
RA (hh mm)	04 18			
Dec (dd.d)	+38.0			
GST range (Europe)	22:00 - 12:00			
GST range (US)	03:00 - 17:00			
GST range (Other)				
Band(s)				
Flux density (Total, Jy)	>10			
Flux density (correlated, mJy)	5000 - 7000			
RMS needed (mJy/beam)	5			
Peak/RMS needed	~500			

- (21) Preferred VLBI session or range of dates for scheduling, and why:  
 Oct 2007
- (22) Dates which are NOT acceptable, and why:
- (23) Attach a self-contained scientific justification, not in excess of 1000 words.  
 Preprints or reprints will not be forwarded to the referees.

Information about the capabilities of the VLBA may be found on the World Wide Web by starting at the NRAO home page, <http://www.nrao.edu>, and selecting the VLBA from "Sites and Telescopes."

A brief summary of the capabilities of the EVN antennas is given in the EVN STATUS TABLE in the EVN USER GUIDE, which may be found at [http://www.evlbi.org/user\\_guide/user\\_guide.html](http://www.evlbi.org/user_guide/user_guide.html).

Please include the full postal addresses for first-time users or for those that have moved (if not contact author).