VLBA PROPOSAL COVERSHEET

DEADLINES: 1st of Feb., June, Oct. EMAIL TO: propsoc@nrao.edu OR MAIL TO: Director NRAO, Edgemont Rd. Charlottesville, VA 22903-2475, USA

rcvd:

(1) Date Prepared: January 15, 2007

(2) Title of Proposal: TOO - Interferometric Radar Observations of Asteroid 2006 VV2

			Students Only				
(3) AUTHORS	INSTITUTION	E-mail		For	Ph.D.		
(Add * for new location)			G/U	Thesis?	Year		
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(4) Related previous or current VLBI proposal(s): BB187							
(5) Contact author for scheduling: Greg Black (6) Telephone: 434-243-8941							
Address: PO Box 400325 Fax: 434-924-3104							
	pt. of Astronomy	101. 101.021.0101					
-	University of Virginia						
Charlottesville, VA 22904-4325							
(7) Scientific Category: \bigcirc astrometry & geodesy \bigcirc galactic \bigcirc extragalactic \bigotimes other: Solar System object							
Rapid Response Science: \bigcirc Known Transient \bigcirc Exploratory \bigotimes Target of Opportunity							
(8) Wavelength(s) requested (those not available on the global network are indicated with a small circle):							
$\bigcirc 90 \mathrm{cm} \bigcirc 50 \mathrm{cm} \circ 30 \mathrm{cm} \bigcirc 21 \mathrm{cm} \bigcirc 18 \mathrm{cm} \bigotimes 13 \mathrm{cm} \bigcirc 6 \mathrm{cm} \circ 5 \mathrm{cm} \bigcirc 3.6 \mathrm{cm} \bigcirc 3.6/13 \mathrm{cm}$							
$\circ 2 \mathrm{cm} \bigcirc 1.3 \mathrm{cm} \bigcirc 7 \mathrm{mm} \circ 3 \mathrm{mm}$							
\bigcirc Global Network standard bands \bigcirc Special frequencies:							
(9) Recording format: O Default continuum setup (VLBA only), 🛇 VLBA/MkIV, O MkIII: Mode							
Bandwidth per BaseBand channel: 0.125 MHz							
Aggregate bit rate: 32 (8 BB channels at 2 MSamples/sec of $\bigcirc 1$ bit, $\bigotimes 2$ bit)							
$(10) \bigcirc$ Multi-epoch obser	vation: epochs of	hours each, separated by					
(11) Network	Requested	antennas	Fotal tin	ne requeste	d		
EVN & MERLIN							
VLBA	ALL	2 hr					
other NRAO	GBT	2 hr					
Non-VLBI Intruments							

(12) ABSTRACT (Do not write outside this space. Please type)

Near Earth asteroid 2006 VV2 will approach the Earth to within 0.024 AU in late March-early April 2007 at which time we request a short VLBA+GBT session to image and do astrometry of this asteroid while it's illuminated by the Arecibo radar. Synthesis radar imaging can provide plane-of-sky image and position information orders of magnitude better than other ground-based methods, and which is complementary to standard radar mapping. The asteroid's diameter is likely between 1.5 and 2 km, and we expect a spatial resolution of order 50 m. We request the raw data streams in order to correlate in software to achieve a narrow frequency resolution matched to the object's bandwidth which depends on its unknown rotation rate but is likely ≤ 20 Hz. However, the echo from this object may be strong enough to be seen at the relatively coarse frequency resolution of the VLBA correlator.

- (13) Observation type: \otimes Interferometry, \otimes Spectroscopy, \bigcirc Pulsar, \otimes Phase referencing
- (14) Proposal is \bigcirc Suitable \bigotimes Unsuitable for dynamic scheduling.
- (15) Polarization: \bigcirc Single Polarization \bigotimes Dual Circular Polarization Global network standard for single polarization is LCP for all λ s except 13cm (RCP) and 3.6cm (RCP).
- (16) Tape usage (Show <recording time>/<total time>):
- (17) Assistance required:
 Observation Setup:
 O Consultation,
 Consultation,
 Extensive help,
 O Observe file preparation
 Consultation,
 Extensive help,
 Calibration service
- (18) Processor: ⊗ Socorro, JIVE, Haystack, Bonn, Washington, Other______ Special processing: ○ XPol, ○ Pulsar gate, ○ Multiple Fields: ______ Averaging time: ______ Spectral channels per baseband channel: 1024 ⊗ Other special processing: need to extract raw data

(19) Postprocessing Location: <u>home</u>

(20) Source list: \bigotimes J2000 \bigcirc B1950

If more than 4 sources, please attach list. If more than 30, give only selection criteria and GST range(s)

	Source 1	Source 2	Source 3	Source 4
Name(s)	Asteroid 2006 VV2			
RA (hh mm)	10 37			
Dec (dd.d)	+06.2			
GST range (Europe)				
GST range (US)	14:00 - 16:00			
GST range (Other)				
Band(s)	13 cm			
Flux density (Total, Jy)				
Flux density (correlated, mJy)				
RMS needed (mJy/beam)				
Peak/RMS needed				

- (21) Preferred VLBI session or range of dates for scheduling, and why: 2007 Apr 1. Must coincide with asteroid close approach and transmit window from Arecibo. 2007 Mar 31 could be an optional date, but less optimal.
- (22) Dates which are NOT acceptable, and why: All others. Object is only observable at requested times.
- (23) Attach a self-contained scientific justification, not in excess of 1000 words. Preprints or reprints will not be forwarded to the referees.

Information about the capabilities of the VLBA may be found on the World Wide Web by starting at the NRAO home page, http://www.nrao.edu, and selecting the VLBA from "Sites and Telescopes."

A brief summary of the capabilities of the EVN antennas is given in the EVN STATUS TABLE in the EVN USER GUIDE, which may be found at http://www.evlbi.org/user_guide/user_guide.html.

Please include the full postal addresses for first-time users or for those that have moved (if not contact author).